

Aerosol Speciation Using UV-Vis Observations by DSCOVR/EPIC

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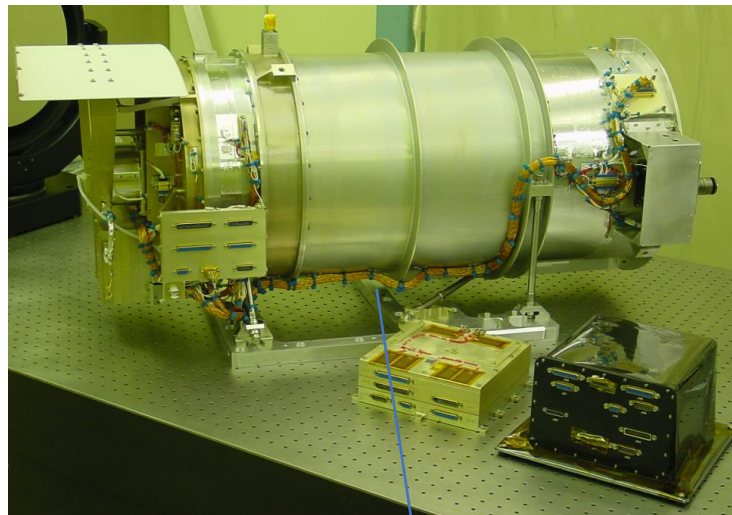


CEOS Atmospheric Composition Virtual Constellation AC-VC-17

June 7-11, 2021

Earth Polychromatic Imaging Camera (EPIC)

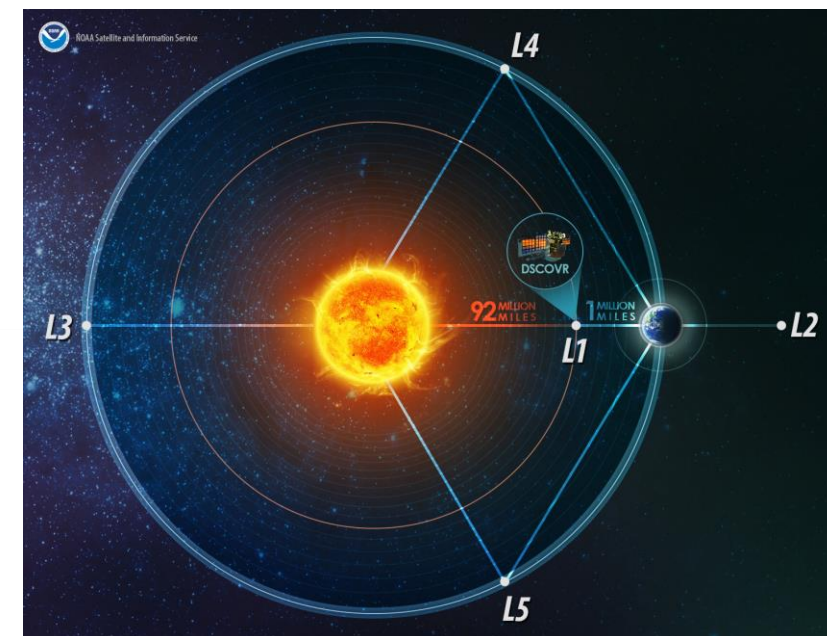
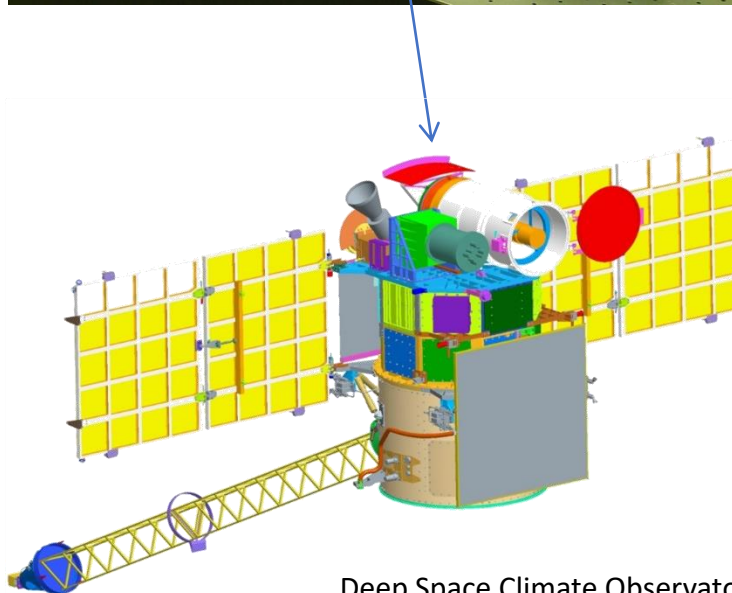
- 2048 x 2048 pixel CCD;
- 8 km pixel; 2x2 onboard aggr.
- Number of daytime images
6 in winter (same area)
Up to 12 in summer



On orbit since 2015

Wavelength (nm) Full width (nm) Primary Applications

317.5 ± 0.1	1 ± 0.2	Ozone, SO_2
325 ± 0.1	2 ± 0.2	Ozone
340 ± 0.3	3 ± 0.6	Ozone, Aerosols
388 ± 0.3	3 ± 0.6	Aerosols, Clouds
443 ± 1	3 ± 0.6	Aerosols
551 ± 1	3 ± 0.6	Aerosols, Vegetation
680 ± 0.2	2 ± 0.4	Aerosol, Vegetation, Clouds
687.75 ± 0.2	0.8 ± 0.2	Cloud Height
764.0 ± 0.2	1 ± 0.2	Cloud Height
779.5 ± 0.3	2 ± 0.4	Clouds, Vegetation

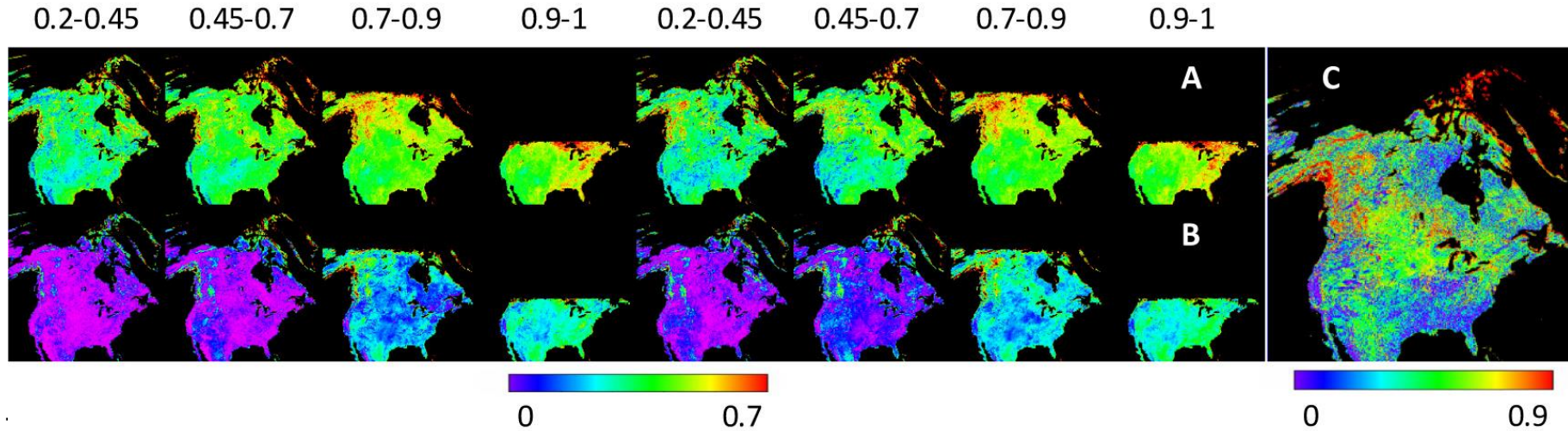


Deep Space Climate Observatory (DSCOVR)

Aerosol Retrieval
Atmospheric Correction

MAIAC EPIC v2 Joint Retrieval Algorithm

- Minimum Reflectance Method to characterize spectral SR ratio at 340, 388, 443 and 680nm in 4 SZA bins



Example for NA:
 (A) 443/680nm
 (B) 388/680nm
 (C) 340/388nm

- Absorption model (imaginary ref. index): $k_\lambda = k_0 (\lambda / \lambda_0)^{-b}$ for $\lambda < \lambda_0$, where $\lambda_0 = 680\text{nm}$
(in the limit of small particles, $AAE \sim b+1$, where Absorption Ångström Exponent AAE is defined for the AAOD).
- Real *refIM* and size distribution are fixed. The results are reported for $H^a=1$ and 4km for smoke and dust.
- Retrieve $\{AOD, k_0, b\}$ using Levenberg-Marquart optimal fit of 340, 388, 443 and 680nm:

$$F^2 = 1/N \sum \left(\frac{L_\lambda^m - L_\lambda^t}{L_\lambda^m} \right)^2 = \min\{AOD_{443}, k_0, b\}$$

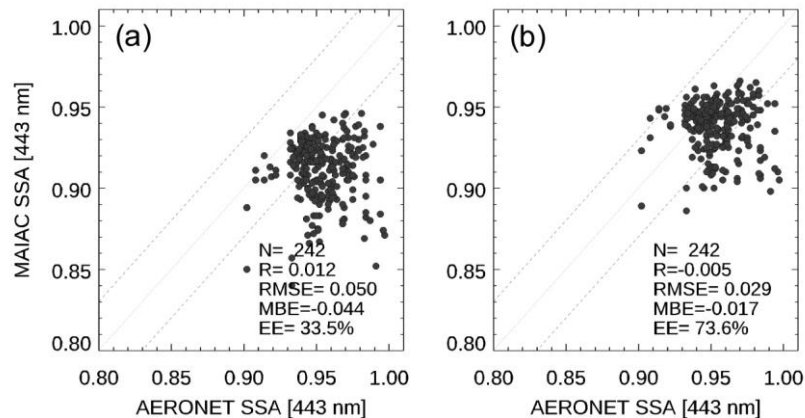
- LUT-based retrievals on a 4x4 matrix of $b = \{0.1, 1.5, 3, 4\}$ and $k_0 = \{0.001, 0.006, 0.011, 0.016\}$ - smoke
 $\{0.0006, 0.0014, 0.0022, 0.003\}$ - dust.

Summary SSA₄₄₃ Validation: 2018

Smoke, North America

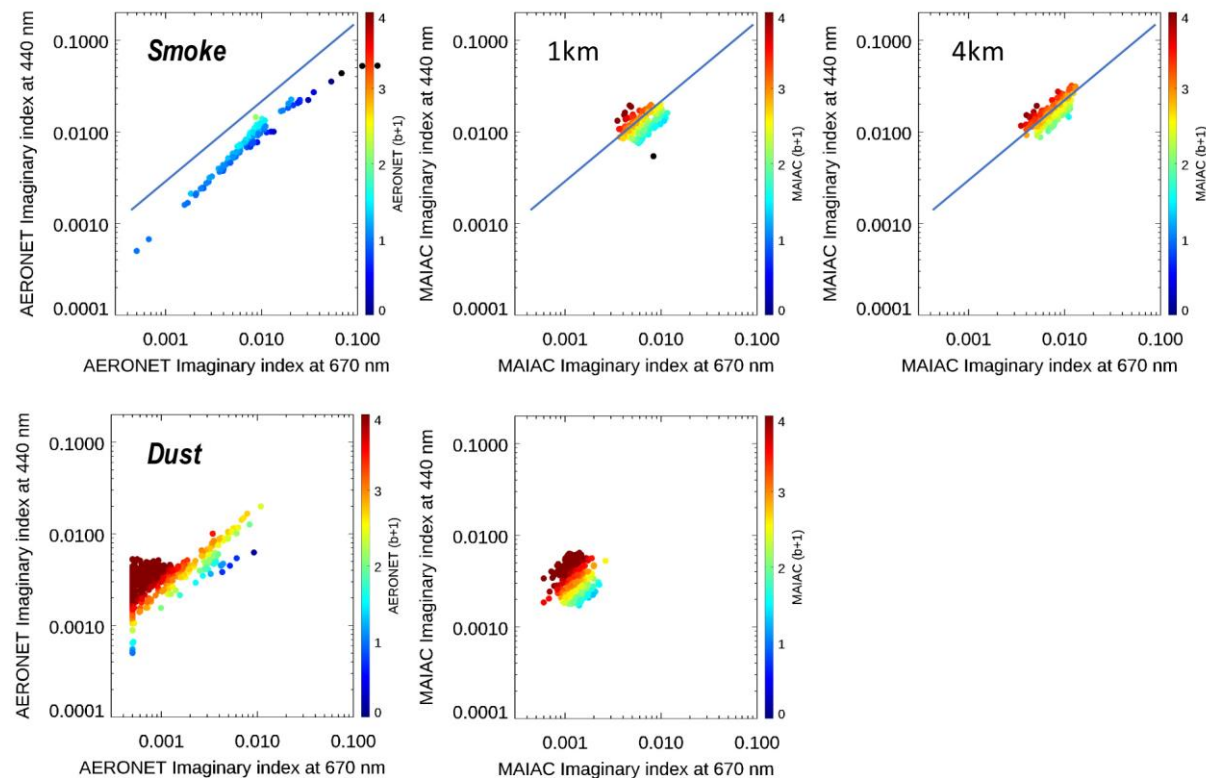
H^a=1km

H^a=4km

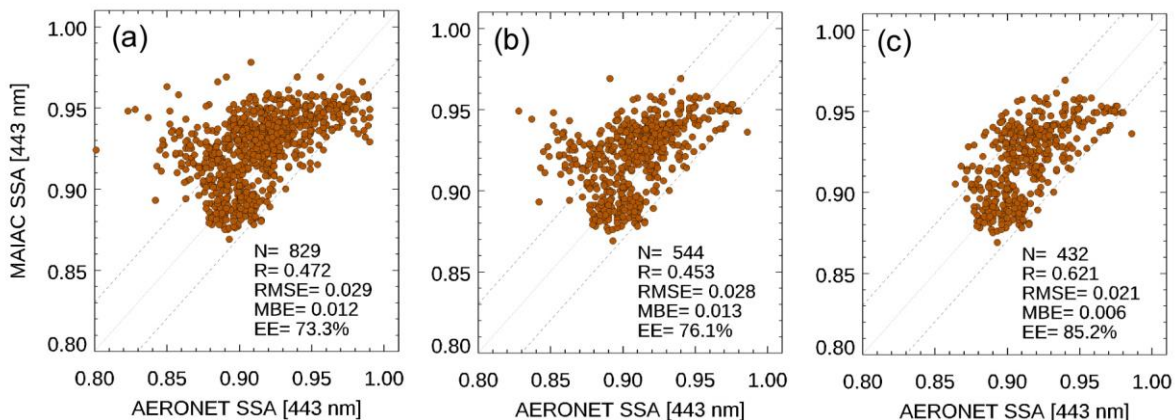


Spectral Dependence of Absorption

Following Figure 3 and Figure 4 of Schuster et al (2016)



Dust: North Africa/Arabia/Middle East



(A,B) 17 sites, AERONET AOD₄₄₀ > 0.4 and > 0.6.

(C) 15 sites (AOD₄₄₀ > 0.6), with Cairo_ENM_2 and Ilorin excluded.

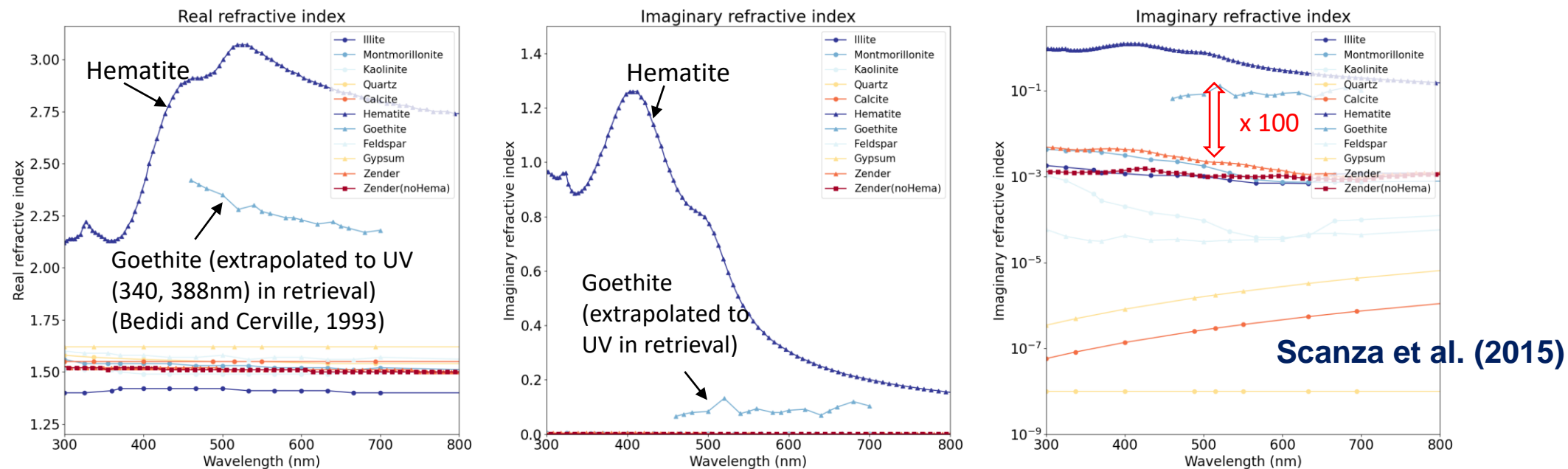
Smoke: MAIAC EPIC *b* is higher than AERONET

(this is expected from UV compared to AERONET's Blue-1020nm)

Dust: EPIC and AERONET *b* are similar

Inferring iron oxide species in mineral dust aerosols

(follow Schuster et al., 2016)



- Hematite ($\alpha\text{-Fe}_2\text{O}_3$) and goethite ($\alpha\text{-FeOOH}$), both in the Fe(III) oxidation state, are the major iron oxides species in mineral dust (Torrent et al., 1983). They are major components controlling the absorption signal magnitude of pure dust toward SW radiation [e.g., Sokolik and Toon, 1999; Moosmüller et al., 2012; Lafon et al., 2006; Formenti et al., 2014], as can be inferred from their complex refractive index imaginary part characteristics (Fig.1).
- We inferred hematite ($\alpha\text{-Fe}_2\text{O}_3$) / goethite ($\alpha\text{-FeOOH}$) content for major global dust source regions from MAIAC EPIC L2 products.

Methodology - (Forward)

(1) With known refractive index (\mathbf{n}, \mathbf{k}) or complex dielectric function (ϵ_r, ϵ_i)

$$\epsilon_1 = \epsilon_{1,r} + i\epsilon_{1,i} = (n_1^2 - k_1^2) + i(2n_1k_1)$$

$$\epsilon_2 = \epsilon_{2,r} + i\epsilon_{2,i} = (n_2^2 - k_2^2) + i(2n_2k_2)$$

$$\epsilon_h = \epsilon_{h,r} + i\epsilon_{h,i} = (n_h^2 - k_h^2) + i(2n_hk_h)$$

, where 1, 2, h indicates inclusion 1 (hematite), inclusion 2 (goethite), and host

(2) Maxwell Garnett effective medium approximation (f_1, f_2 : volume fraction of inclusions)

$$\epsilon_{MG} = \epsilon_h \left[1 + \frac{3 \left(f_1 \frac{\epsilon_1 - \epsilon_h}{\epsilon_1 + 2\epsilon_h} + f_2 \frac{\epsilon_2 - \epsilon_h}{\epsilon_2 + 2\epsilon_h} \right)}{1 - f_1 \frac{\epsilon_1 - \epsilon_h}{\epsilon_1 + 2\epsilon_h} - f_2 \frac{\epsilon_2 - \epsilon_h}{\epsilon_2 + 2\epsilon_h}} \right] = \epsilon_{MG,r} + i\epsilon_{MG,i}$$

$$n_{mix} = \sqrt{\frac{\sqrt{\epsilon_{MG,r}^2 + \epsilon_{MG,i}^2} + \epsilon_{MG,r}}{2}}, \quad k_{mix} = \sqrt{\frac{\sqrt{\epsilon_{MG,r}^2 + \epsilon_{MG,i}^2} - \epsilon_{MG,r}}{2}}$$

$$\Rightarrow m_{mix}(\lambda_j) = F(f_1, f_2, m_1(\lambda_j), m_2(\lambda_j), n_{host}(\lambda_j)) \\ = n_{mix}(\lambda_j) + ik_{mix}(\lambda_j)$$

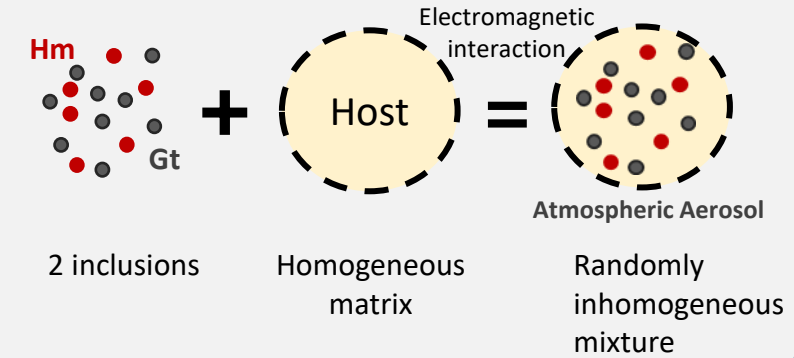
- (Inversion)

$$\chi^2 = \sum_{j=1}^4 \left[\frac{(k_{epic}(\lambda_j) - k_{mix}(\lambda_j))^2}{k_{epic}(\lambda_j)} \right] \rightarrow \min$$

MAIAC EPIC refractive index ($k_{epic}(\lambda_j) = k_{680} (\lambda_j/\lambda_{680})^{-b}$)
Updated iteratively

Illustration of Maxwell Garnett effective medium approximation (**pure dust**)

(Bohren and Huffman, 1987; Schuster et al., 2016)



- **Assumption:** major absorption caused by hematite and goethite (not host)
- $n_h = 1.52$ (fixed), $k_h = 0.0$
- λ_j : 340, 388, 443, 680 nm
- **Fitting k_{epic} part only**, because MAIAC EPIC does not retrieve real refractive index (n_{epic})

From EPIC AOD to composition mass concentration

Volume concentration ($\mu\text{m}^3/\mu\text{m}^2$):

$$C_V = \int_{r_{\min}}^{r_{\max}} \frac{dV(r)}{d \ln r} d \ln r.$$

- Step 1: calculate total (fine + coarse) volume concentration of dust (440 nm)
 - $\tau^a = \tau_f^a + \tau_c^a = C_{Vf}h_f + C_{Vc}h_c \approx C_{Vc}h_c$ ($\because C_{Vf} \ll C_{Vc}$ for AOD>0.6)
 - With $h_c = 1.2526$, $C_V \approx C_{Vc} = \frac{\tau^a}{1.2526} = 0.7983 * \tau^a$
 - C_V : volume concentration ($\mu\text{m}^3/\mu\text{m}^2$)
 - h_c : AOD per unit volume concentration
- Step 2: separate Hematite/Goethite volume concentration (C_V) using retrieved volume fraction (f) results
 - $C_{V,\text{hematite}} = C_V * f_{\text{hematite}}$
 - $C_{V,\text{goethite}} = C_V * f_{\text{goethite}}$
 - $C_{V,\text{host}} = C_V * f_{\text{host}} = C_V * (1 - f_{\text{hematite}} - f_{\text{goethite}})$
- Step 3: Then, calculate Hematite/Goethite mass concentration (C_M)
 - $C_{M,\text{Hematite}} = C_{V,\text{Hematite}} * \rho_{\text{Hematite}}$ (ρ_{Hematite} = mass concentration per unit volume (or density) = 5260 kg/m³)
 - $C_{M,\text{Goethite}} = C_{V,\text{Goethite}} * \rho_{\text{Goethite}}$ (ρ_{Goethite} = 3800 kg/m³)
 - $C_{M,\text{Host}} = C_{V,\text{Host}} * \rho_{\text{Host}}$ ($\rho_{\text{Host}} = \rho_{\text{Zender}}^* = 2500 \text{ kg/m}^3$)
 - * Zender (Mahowald et. al., 2006) assuming Maxwell–Garnett mixing of 47.6 % quartz, 25 % illite, 25 % montmorillonite, 2 % calcite and 0.4 % hematite by volume with density equal to 2500 kg/m³ and hygroscopicity prescribed at 0.14 (Scanza et al., 2015).
 - * Density of free iron is roughly twice that of other minerals (Schuster et al., 2016; Formenti et al., 2014).
- ex) $\tau^a=0.75$, $f_{\text{hematite}}= 0.005$

$$C_{M,\text{Hematite}} = \frac{\tau^a}{1.2526} * f_{\text{hematite}} * \rho_{\text{Hematite}} = \frac{0.75}{1.2526} * 0.005 * 5260 = 15.74 \text{ [mg/m}^2\text{]}$$

Dust episodes - Sahara / Sahel

TOA RGB
(0-0.55)

AOD₄₄₃
(0-2)

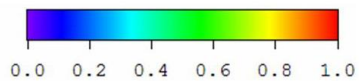
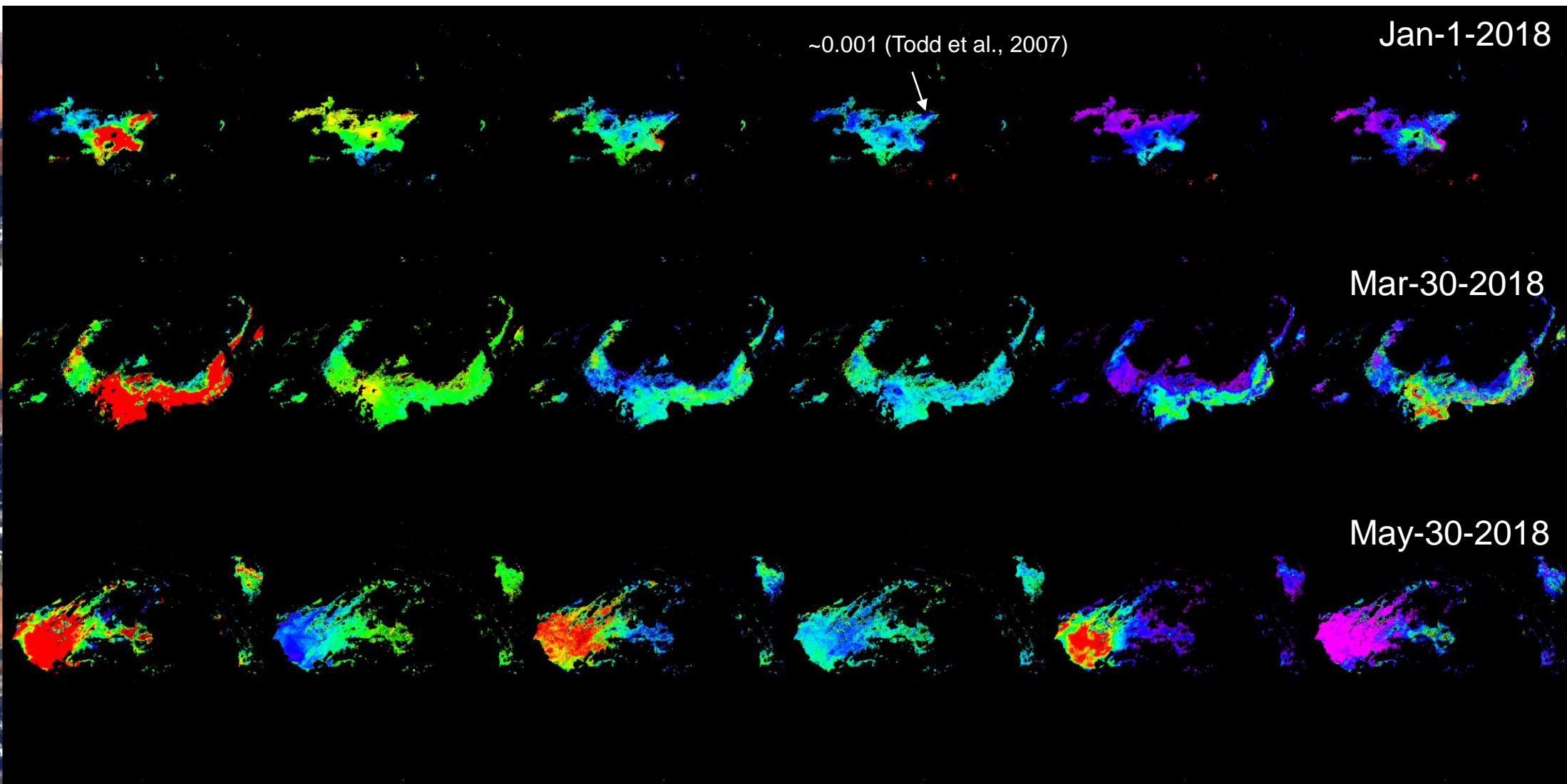
SSA₄₄₃
(0.85-0.98)

b
(0-4)

k_0
(0-0.003)

$C_{M, hematite}$
(0-150) [mg/m²]

$C_{M, goethite}$
(0-150) [mg/m²]



Jan-1-2018

Mar-30-2018

May-30-2018

~0.001 (Todd et al., 2007)

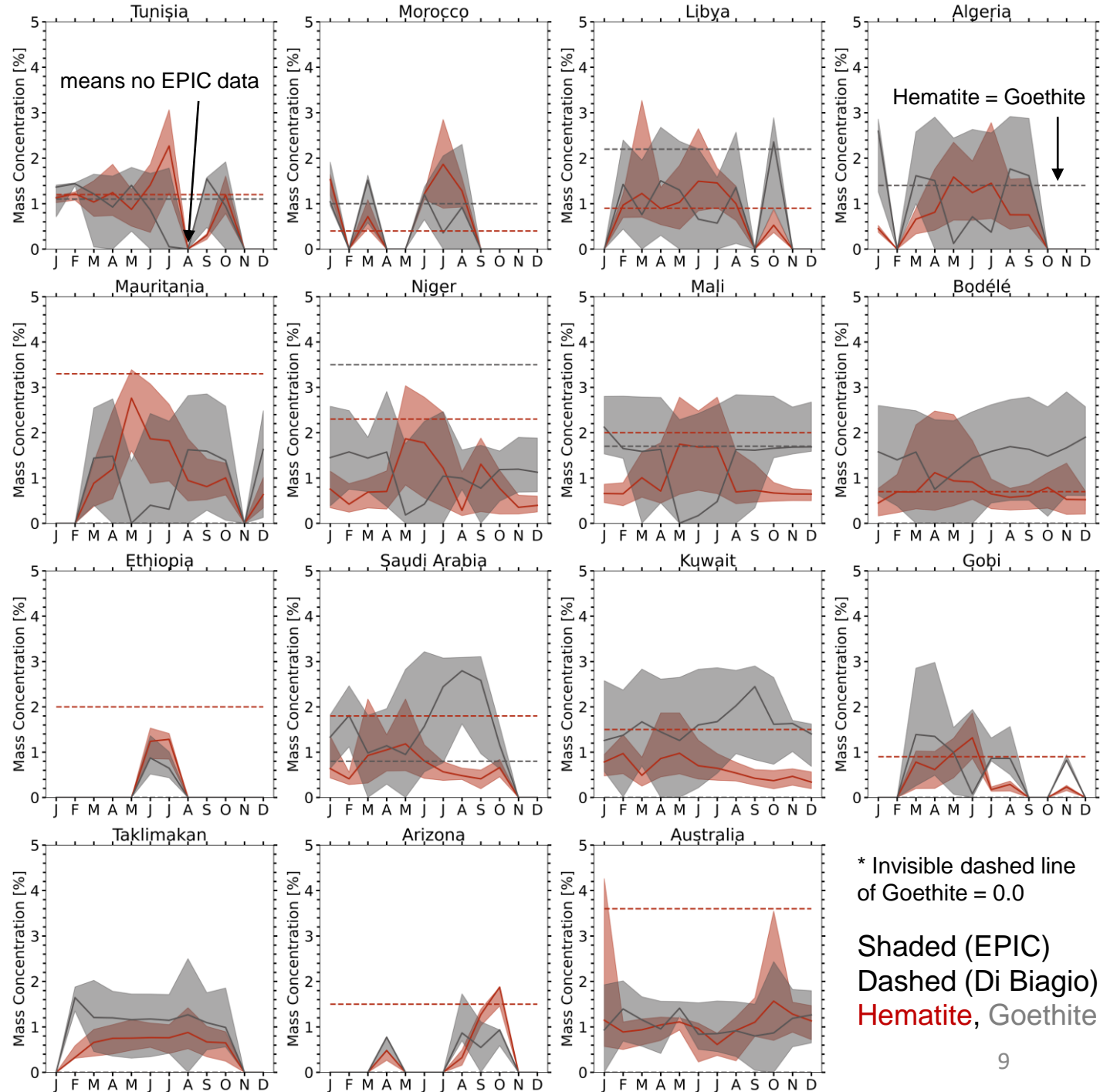
Comparison with Di Biagio et al. (2019)

- **Shaded area : EPIC retrieved**

- Hematite / Goethite
- 5th, median, 95th
- With pixels of AOD>1.0 used only
- ± 1 degree box pixels collected (Monthly)
- 01/01/2018 – 12/31/2018 (1 year)
- **Soil content + affected by transport due to different source regions**

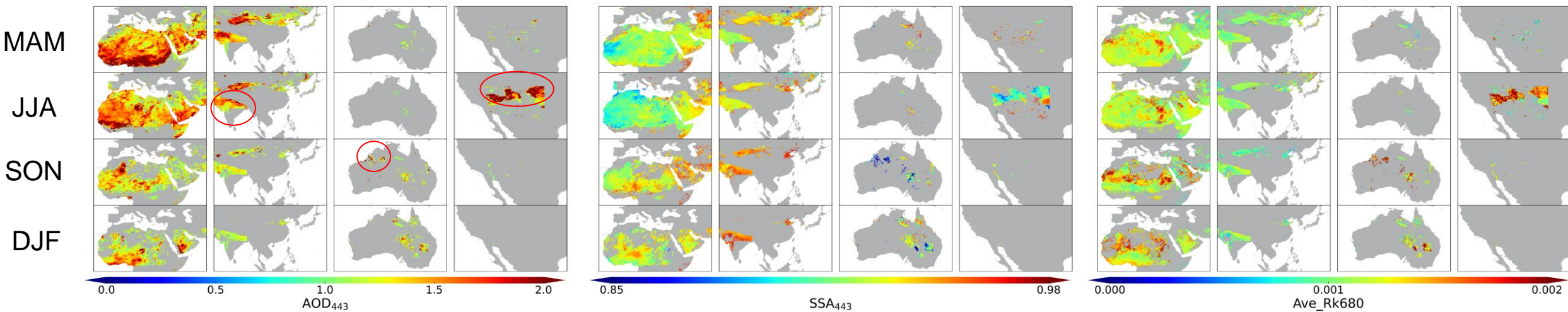
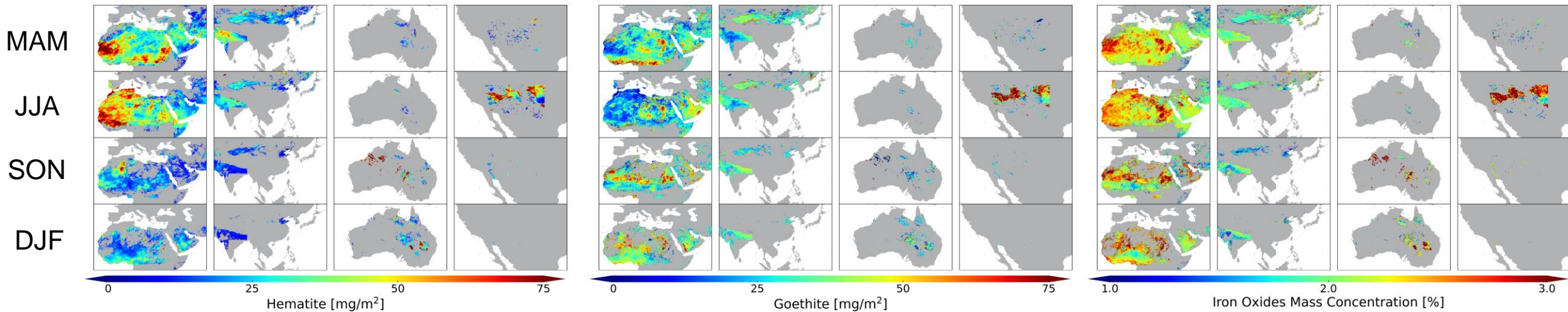
- **Dashed line : Di Biagio et al. (2019)**

- ± 10% uncertainty
- Simulation chamber study (with X-ray absorption near edge structure (XANES) method), from soil samples and sediments collected from each desert area
- Refer to the [bulk](#) composition of [pure dust aerosol](#) in [dry condition](#) with a [size range of 2-6day](#) transport.
- **Soil content only**

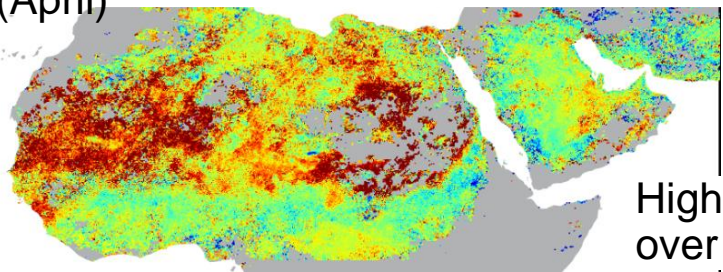


Climatology of iron oxides species

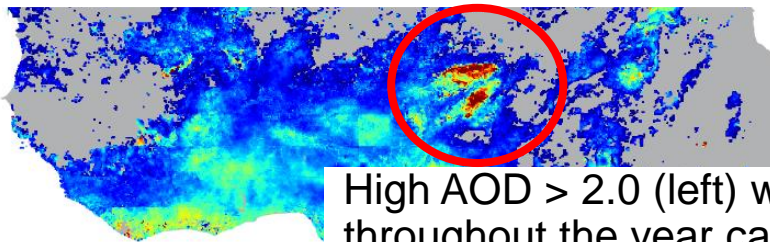
- Pixels of AOD>1.0 used only
- 01/01/2018 – 12/31/2018 (1 year)



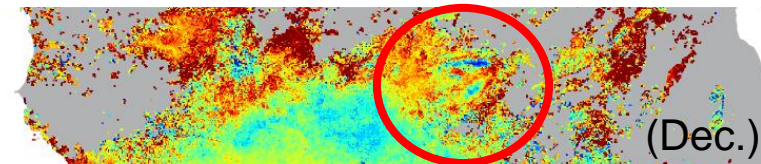
(April)



High Iron oxide over Sahel



High AOD > 2.0 (left) with low iron oxides content ~1% (right) throughout the year captured over Bodélé.



(Dec.)

Conclusions

- **MAIAC EPIC DSCOVR v2 features:**

- New joint retrieval of AOD and spectral absorption
- Good SSA accuracy, $R \sim 0.62$, $rmse \sim 0.02$, bias ~ 0.006 (dust) and $rmse \sim 0.029$, $EE = 73.6\%$ (smoke @ 4km) (AERONET SSA uncertainty ± 0.03)
- Resolve spatial variations of absorption, in particular for Spectral Absorption Exponent (SAE, b)

Lyapustin A., Go S., Korokin S., Wang Y., Torres O., Jethva H. and Marshak A. (2021) Retrievals of Aerosol Optical Depth and Spectral Absorption From DSCOVR EPIC. *Front. Remote Sens.* 2:645794. doi: 10.3389/frsen.2021.645794.

- **Global Analysis of Hematite/Goethite in Mineral Dust Based on (AOD, k_0 , b):**

- Retrieved iron oxides enveloped the overall range of Di Biagio et al. (2019) soil measurement data of iron oxides 0.7-5.8% and were in line with the previous published results
- Hematite/Goethite ratio shows significant seasonal and spatial variability
- Initial Climatology of Hematite and Goethite content is provided

Go, S., A. Lyapustin, G. Schuster et al. Inferring iron oxides content in mineral dust aerosols from DSCOVR EPIC (to be submitted).

- **Similar Analysis of Black/Brown Carbon for Biomass Burning Aerosol is Coming Soon**